Progress and Challenges in Computational Geodynamics

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Lecture III: Magma Dynamics

Outline

Overview and Motivation Equations for coupled fluid-solid dynamics Basic Physics: Localization phenomena on-linear waves Shear band formation, reactive flow Geodynamics Applications : Mid Ocean Ridges and Subduction Zones Open Questions/Future Directions © 2010 Europa Technologies

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- Goals of this lecture
 - develop better physical intuition into basic physics of magma dynamics
 - understand the motivation for developing better abstractions
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 for multi-physics solvers
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- Mantle convection = Convection with Plates
- Plates are defined by their weak boundaries.
- Convergent and Divergent Boundaries are fundamentally magmatic
- How does magmatism affect the dynamics and structure of plate boundaries and global mantle convection?

Why Magma Dynamics? Global Geochemical Evolution



Brandenburg et al, EPSL 2008, 2-D Cylindrical High Ra convection calculation
Solid State Convection primarily stirs
Chemical Fractionation, mixing and sampling of the mantle requires a mobile liquid phase

• Can we use variation in composition of erupted lavas to infer rate and efficiency of convecting stirring in Earth?



- Large Scale Deformation of the Earth is in the solid-state
- Most melting occurs in small scale regions near plate boundaries, but may affect global flow and plate tectonics
- How do we understand the basic physics and interactions across scales and constrain it with chemical data?



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- In the absence of liquids, the system must be consistent with mantle convection (viscously deformable)

(McKenzie, 1984, JPet; Scott & Stevenson, 1984, 1986, JGR; Bercovici, Ricard et al., 2001, 2003; Simpson et al, 2010 JGR)

Conservation of Mass: Fluid

$$\frac{\partial(\rho_f \phi)}{\partial t} + \boldsymbol{\nabla} \cdot (\rho_f \phi \mathbf{v}) = \Gamma$$

Conservation of Mass: Solid

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Plus Constitutive Relations/Closures

Permeability	$K \sim k_0(d, \ldots)\phi^n$
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Coupled through pressure

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A Better (?) Formulation (<u>McKenzie Tutorial Notes</u> @ CIG, Katz et al, 2007 Pepi)

$$\boldsymbol{\nabla} P = \bar{\rho} \mathbf{g} + \boldsymbol{\nabla} \left(\zeta - \frac{2\eta}{3} \right) \boldsymbol{\nabla} \cdot \mathbf{V} + \boldsymbol{\nabla} \cdot \eta \left(\boldsymbol{\nabla} \mathbf{V} + \boldsymbol{\nabla} \mathbf{V}^T \right)$$

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Decompose the pressure into 3 terms

$$P = P_l + \mathcal{P} + P^*$$

with

• Lithostatic Pressure,
$$P_l = \rho_s^0 g z$$

- "Compaction Pressure", $\mathcal{P} = (\zeta 2\eta/3) \boldsymbol{\nabla} \cdot \mathbf{V}$
- Dynamic Pressure, P^*

A Better (?) Formulation (<u>McKenzie Tutorial Notes</u> (CIG), Katz et al, 2007 Pepi)

$$\begin{split} \frac{D\phi}{Dt} &= (1-\phi)\frac{\mathcal{P}}{\xi} + \Gamma/\rho_s \quad \begin{array}{l} \text{Compressible} \\ \text{Flow} \end{array} \\ &- \nabla \cdot \frac{K}{\mu} \nabla \mathcal{P} + \frac{\mathcal{P}}{\xi} = \nabla \cdot \frac{K}{\mu} \left[\nabla P^* + \Delta \rho \mathbf{g} \right] + \Gamma \frac{\Delta \rho}{\rho_f \rho_s} \\ &\nabla \cdot \mathbf{v} = \frac{\mathcal{P}}{\xi} \quad \begin{array}{l} \text{``Incompressible''} \\ \text{Flow} \end{array} \\ &\nabla P^* = \nabla \cdot \eta \left(\nabla \mathbf{v} + \nabla \mathbf{v}^T \right) - \phi \Delta \rho \mathbf{g} \end{split}$$

with

•
$$\xi = (\zeta - 2\eta/3) = \eta \left(\frac{1}{\phi} - \frac{2}{3}\right) \approx \eta/\phi$$

•
$$\Delta \rho = \rho_s - \rho_f$$

Comparison to Thermal Convection

(McKenzie Tutorial Notes (CIG/bSpace), Katz et al, 2007 Pepi)

$$\frac{D\phi}{Dt} = (1 - \phi)\frac{\mathcal{P}}{\xi} + \Gamma/\rho_s \quad \text{``Magma''}$$
$$-\nabla \cdot \frac{K}{\mu}\nabla \mathcal{P} + \frac{\mathcal{P}}{\xi} = \nabla \cdot \frac{K}{\mu} [\nabla P^* + \Delta \rho \mathbf{g}] + \Gamma \frac{\Delta \rho}{\rho_f \rho_s}$$
$$\nabla \cdot \mathbf{v} = \frac{\mathcal{P}}{\xi}$$
$$\nabla P^* = \nabla \cdot \eta (\nabla \mathbf{v} + \nabla \mathbf{v}^T) - \phi \Delta \rho \mathbf{g}$$
Eq.
$$\frac{DT}{Dt} = \nabla^2 T \qquad \text{Thermal Convection}$$
$$\nabla \cdot \mathbf{v} = 0$$
$$\nabla P = \nabla \cdot \eta (\nabla \mathbf{v} + \nabla \mathbf{v}^T) - \text{Ra}T\mathbf{g}$$

Non-linear wave equations for porosity (Scott & Stevenson, Nature, 1984, Spiegelman, JFM 1993, Simpson & Spiegelman, JSC 2010)

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Intrinsic length Scale: The compaction length

(McKenzie, JPet, 1984, Scott & Stevenson, Nature, 1984, Spiegelman, JFM 1993)

The Compaction Length

$$\delta = \sqrt{\frac{K(\phi)\zeta(\phi)}{\mu}}$$

PermeabilityKSolid Bulk Viscosity ζ melt Shear Viscosity μ

$${\cal K}(\phi)\propto \phi^n$$

 ${\cal L}(\phi)\propto \eta/\phi^m$
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Permeability Solid Bulk Viscosity $\zeta(\phi) \propto \eta/\phi^m$ melt Shear Viscosity

 $\mathsf{K}(\phi) \propto \phi^{\mathsf{n}}$ μ

Length scale of pressure variations due to a change in flux.



(Scott & Stevenson, 1984 Nature, Spiegelman, JFM, 1993, Simpson et al 2008)



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•Variations in melt flux propagate as *non-linear* porosity waves

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Variations in melt flux propagate as *non-linear* porosity waves
Speed and structure of porosity waves depends on *permeability* and *solid rheology*

(Scott & Stevenson, 1984 Nature, Spiegelman, JFM, 1993, Simpson et al 2009,2010)



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•Collision of 2, 2D-porosity waves. P2-P2 FEM with Semi-Lagrangian 2nd-order time stepping. Hybrid FEniCS/PETSc codes.

(Scott & Stevenson, 1984 Nature, Spiegelman, JFM, 1993, Simpson et al 2009,2010)



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 Instability of ID- 3-D waves. 3-D mixed finite elements. Hybrid FEniCS/PETSc codes. (CIG)

(Scott & Stevenson, 1984 Nature, Spiegelman, JFM, 1993, Simpson et al 2009,2010)



(Scott & Stevenson, 1984 Nature, Spiegelman, JFM, 1993, Simpson et al 2009,2010)



Instability of ID- 3-D waves. Spiegelman and Wiggins, 1994, GRL.
 FV geometric multi-grid code.

•Wave behavior is the natural consequence of non-linearity of flux with porosity and viscous deformation of the solid.

- •Waves are generated by obstructions in the flux.
- Implies that magma dynamics is highly time dependent

Solitary waves provide an excellent nonlinear benchmark for space-time codes.
Simpson and Spiegelman, JSC, 2010 provides sinc-codes for calculating spectrally accurate wave profiles in 1, 2 and 3-D.





Full equations with porosity weakening shear viscosity η(φ, V)
Neglect gravity (at lab scale)
PETSc codes with segregated SNES
Spontaneously develops shear band instability







Other sources of melt channelization

Reactive infiltration instability



Reactive Infiltration Instability



Reactive Infilitration Instability



Reactive Infiltration Instability



Chem al Consequences of Melt Channeling

(Spiegelman & Kelemen, 2003, G3)



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- length scale of features controlled by the compaction length (0-10km)

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- What are the observable consequences of these mechanisms.
- How/do small scale physics influence large scale mantle dynamics?

Petascale AMR FEM/Rhea



Global Convection code with parallel adaptive mesh refinement • minimum mesh spacing ~ I km resolves weak boundaries

Adaptive refinement in weak/ plastic regions •Full refinement at h=1km ~ 10¹² elements (exascale?) •Can accomplish, goal oriented adaptation to convergence with 150-300 million elements (10³-10⁴) savings
Geo problems: mid-ocean ridge models (Courtesy Richard Katz)



Melt and solid flow field for a heterogeneous melting mantle beneath a midocean ridge

Full solution of magma dynamics using the "enthalpy method" Katz, J. Pet, 2008

PETSc parallel, structured FV code on staggered mesh.

Location of Volcanoes in Subduction Zones



Geo problems: Subduction Zone models (Spiegelman, van Keken, Hacker, 2009)



Monday, January 10, 2011

Nicaragua Model

(Syracuse et al, 2009)



Slab H₂0 Model

(B Hacker, Perple_X)



Slab H₂0 Model

(B Hacker, Perple_X)

XH2O 0.04 0.03 0.02 0.01

Nicaragua: Slab Water model Hacker Perple_X Wt % water bound in slab minerals

Z_X

Permeable Flow model on subdomain Spiegelman (MADDs-FP -- CIG)



Monday, January 10, 2011

Fluid Flow Trajectories given dehydration rates



Monday, January 10, 2011

Comparison to TUCAN Data



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- addition of a melt phase introduces new dynamics and new length & time scales
- Many different mechanisms suggest some form of mesoscale organization into melt channels in the mantle which may have significant observational consequences.
- Small changes in couplings can significantly change the physics and computational requirements of these problems

Open Questions

- What are the interactions/dominant mechanism for localization at the mesoscale?
- What are the interaction between mesoscale and plate-boundary scale flow?
 Plate boundary dynamics and global mantle convection?
- What are the observational consequences of these processes and can important inferences be made from existing data on the structure and processes of partially molten regions?

Computational & Software issues

- Magma Dynamics is fundamentally a coupled multiscale, multi-physics problem.
- How do we develop flexible, high-performance tools for more readily exploring the space of models and behavior?
- This is a completely different issue than finding/tuning a well understood problem (eg. Navier-Stokes, Seismic Wave tomography).
- Much of the essential software already exists (e.g. PETSc, FEniCS). Next time will detail how we can use it to develop some flexible and general approach to solving multi-physics models.

Philosophy of multi-physics PDE based models

- Overall Structure and Choices
- Software design for managing choices (PETSc, FEniCS)
- General abstractions of Non-linear multi-physics problem
- Examples in Hybrid FEniCS/PETSc codes.
- HPC issues...