

A Parameter Survey of Black Hole Accretion in Radiation GRMHD Simulations

Radiation and magnetic fields play crucial roles in shaping black hole accretion, particularly in near- and super-Eddington regimes. To model these systems, we solve the GRMHD equations coupled with angle-dependent radiation transfer, which enables us to capture the complex dynamics driven by radiation and magnetic fields in extreme environments. In the super-Eddington regime, radiative support causes the accretion disk to thermally expand, forming a narrow conical funnel through which radiation escapes, leading to low radiation efficiency. In the near- and sub-Eddington regime, the magnetic field topology strongly influences the resulting disk structure, allowing the system to reach a steady state as either a thin disk with magnetic coronae or a magnetically elevated disk. These simulations broadly align with observational findings — such as the soft states of X-ray binaries, ultraluminous X-ray sources, and “little red dots” — and provide predictive diagnostics for future observations, which I will discuss in detail during the talk.



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2:30 - 3:30 p.m.

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