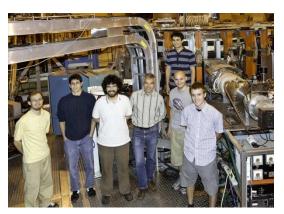
Large-scale electron acceleration by parallel electric fields during magnetic reconnection







J Egedal, A Le, J Ng, O Ohia, A Vrublevskis, P Montag W Daughton & VS Lukin

MIT, PSFC, Cambridge, MA



Fluid simulations





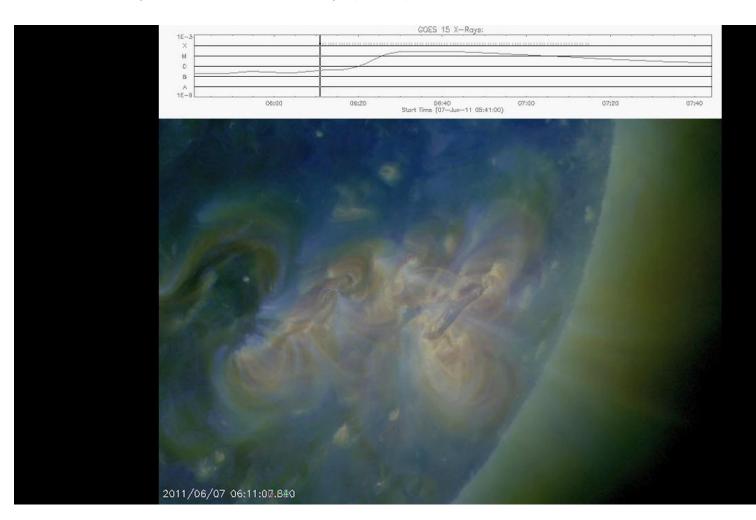


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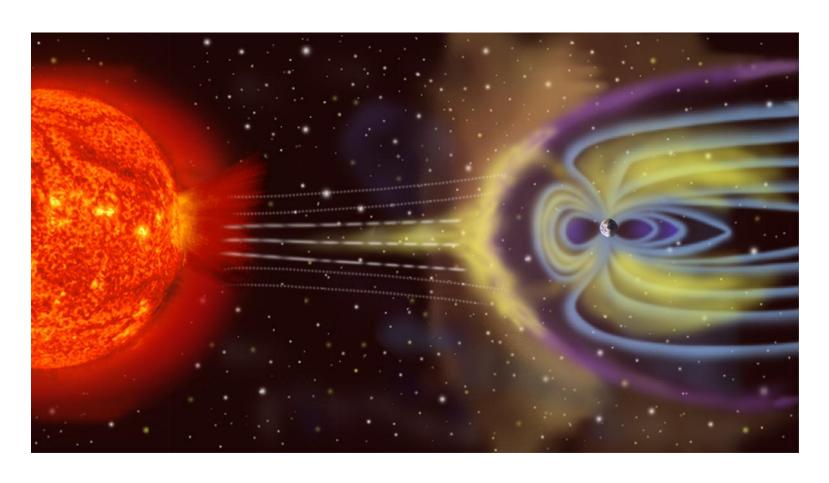
Coronal Mass Ejections

Movie from NASA's Solar Dynamics Observatory (SDO)

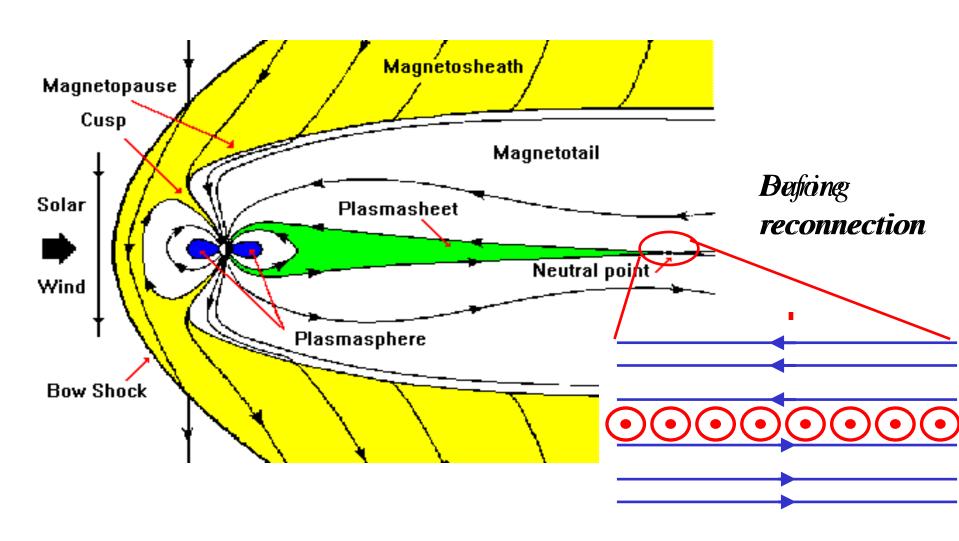


Space Weather

The Solar Wind affects the Earth's environment



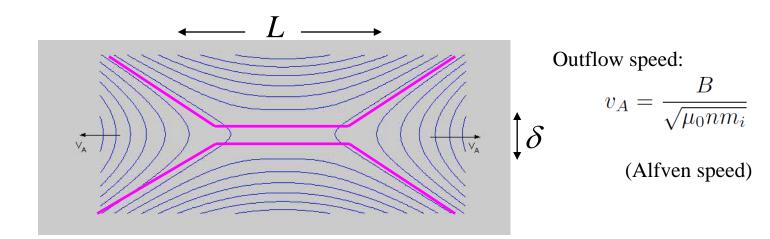
The Earth's Magnetic Shield



Reconnection: A Long Standing Problem

Simplest model for reconnection:

$$\mathbf{E} + \mathbf{v} \times \mathbf{B} = \eta \mathbf{j}$$
 [Sweet-Parker (1957)]



Sweet-Parker: $L >> \delta$:

$$t_{sp} = \sqrt{t_R t_A} = \sqrt{\frac{\mu_0 L^2}{\eta}} \sqrt{\frac{L}{v_A}}$$

Unfavorable for fast reconnection

Two months for a coronal mass ejections

Plasma Kinetic Description

The collisionless Vlasov equation:

$$\left(\frac{\partial}{\partial t} + \mathbf{v} \left(\frac{d}{dt} f_j(\mathbf{x}, \mathbf{v}, t) = 0\right) \cdot \nabla_v\right) f_j = 0$$

$$n_j = \int f_j d^3 v \qquad \qquad \mathbf{J}_j = q_j \int \mathbf{v} f_j d^3 v$$

+ Maxwell's eqs.

Vlasov-Maxwell system of equations

Can be solved numerically (PIC-codes)

Fluid Formulation (Conservation Laws)

mass:
$$\frac{\partial n}{\partial t} + \frac{\partial (nu_j)}{\partial x_j} = 0,$$
momentum:
$$mn\left(\frac{\partial u_j}{\partial t} + u_k \frac{\partial u_j}{\partial x_k}\right) + \frac{\partial P_{jk}}{\partial x_k} - en(E_j + \epsilon_{jkl}u_k B_l) - F_j^{\text{coll}} = 0,$$

energy:
$$\frac{\partial \mathsf{P}_{jk}}{\partial t} + \frac{\partial}{\partial x_l} (\mathsf{P}_{jk} u_l + \mathsf{Q}_{jkl}) + \frac{\partial u_{[j}}{\partial x_l} \mathsf{P}_{lk]} - \frac{e}{m} \epsilon_{[jlm} B_m \mathsf{P}_{lk]} - \mathsf{G}_{jk}^{\mathrm{coll}} = 0$$

Isotropic (scalar) pressure is the standard closure!

$$p = nT$$

Add Maxwell's eqs to complete the fluid model



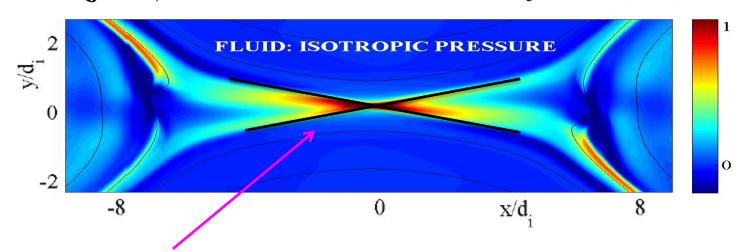
Two-Fluid Simulation

GEM challenge (Hall reconnection)

Out of plane current

$$\mathbf{E} + \mathbf{v} \times \mathbf{B} = (\mathbf{j} \times \mathbf{B})/\text{ne}$$
 [Birn,..., Drake,... Bhattacharjee, et al. (2001)]

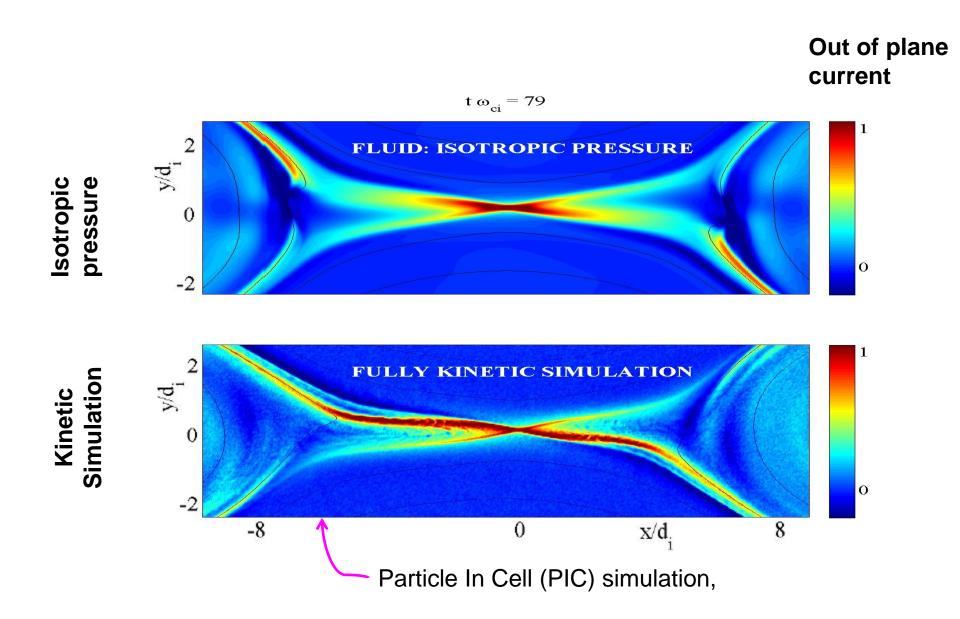
Isotropic pressure



Aspect ratio: 1 / 10

$$\rightarrow$$
 $v_{in} \sim v_A / 10$

Two-Fluid vs Kinetic Simulations



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Wind Spacecraft Observations in Distant Magnetotail, 60*R*_E

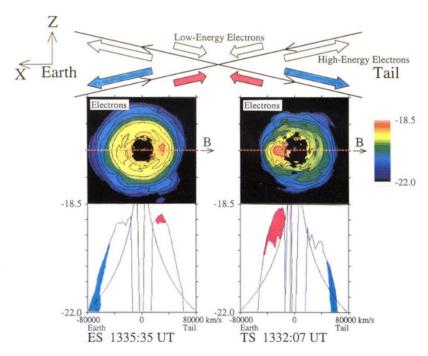
• Measurements within the ion diffusion region reveal:

Strong anisotropy in f_e Wind, $\log_{10}(f/(s^3/km^6))$ $p_{\parallel} > p_{\perp}$

[Øieroset et al., PRL (2001)]

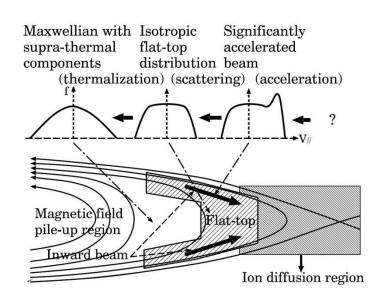
Observed Electron Heating

1-10 keV beam like electron distributions often observed



T. Nagai, et al., *J. Geophys. Res.*, **106**, 25.929 (2001).

Flat-top distributions typical in the exhaust; related beams?

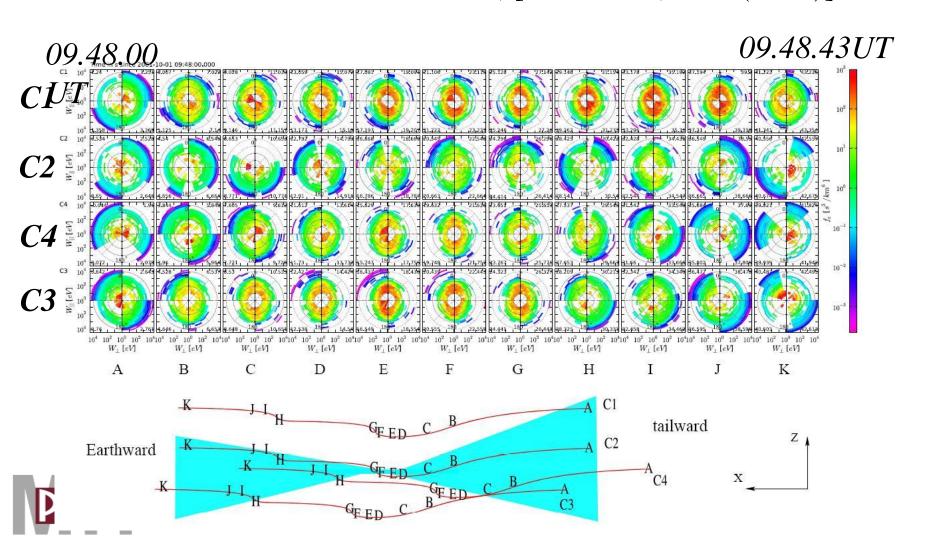


Y. Asano, et al., *J. Geophys. Res.*, *113*(A1), (2008).



Cluster observations

Cluster observations on 2001-10-01, [L.-J. Chen, JGR (2008)]



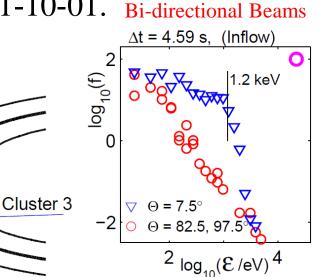
Observed Electron Heating

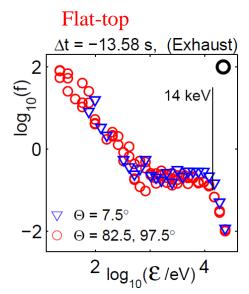
Cluster observations on 2001-10-01. B

Exhaust

Diffusion region

Inflow





Reproduced from Egedal, ..., Chen et al. JGR (2010) Similar figures given in L.-J. Chen, et al JGR (2008), POP (2009)

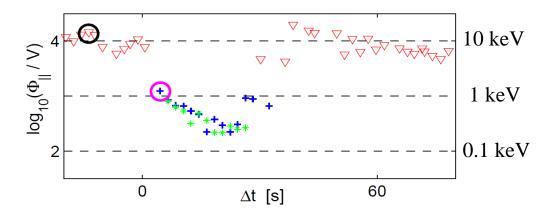
Lobe: $T_e \sim 0.1 \text{ keV}$

Inflow: $T_{e\perp} \sim 0.1 \text{ keV}$, $T_{e\parallel} \sim 1 \text{ keV}$

Exhaust: $T_{e\perp} \sim T_{e\parallel} \sim 10 \text{keV}$



Separator



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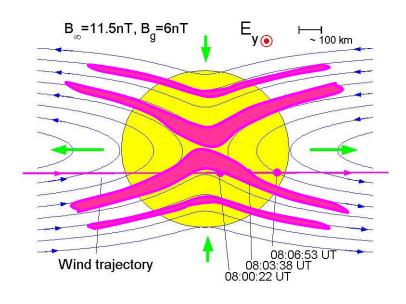


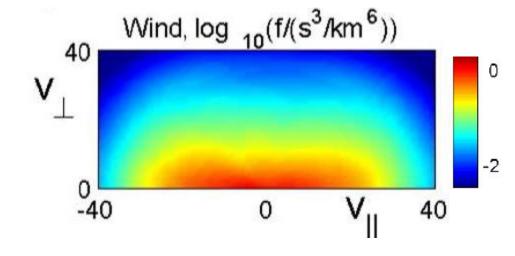
Wind Spacecraft Observations in Distant Magnetotail, 60*R*_E

• Measurements within the ion diffusion region reveal:

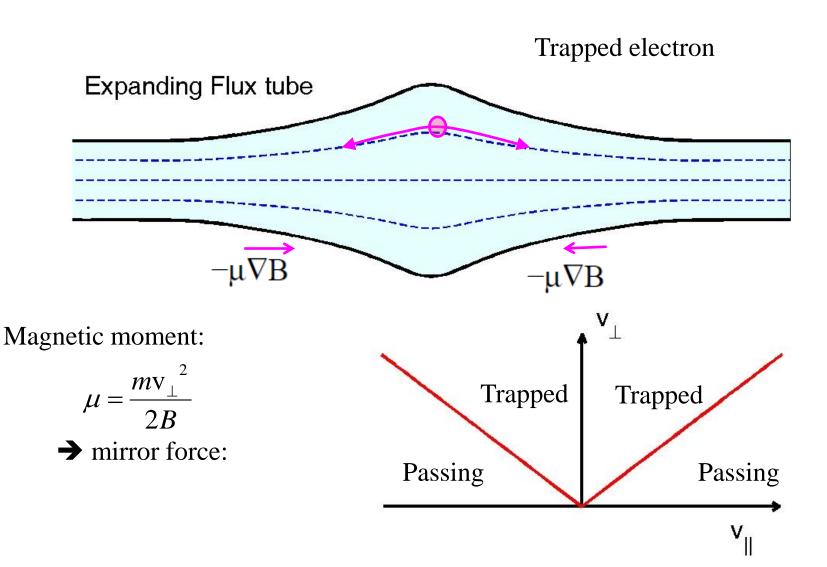
Strong anisotropy in $f_{\rm e}$

$$p_{\parallel} > p_{\perp}$$

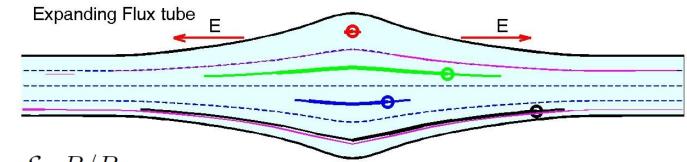




Electrons in an Expanding Flux Tube



Electrons in an Expanding Flux Tube

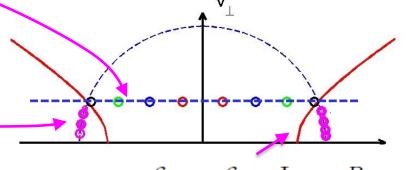


Trapped:

$$\mathcal{E}_{\perp} = \mu B = \mathcal{E}_{\infty} B / B_{\infty} \longrightarrow \mathcal{E}_{\infty} = \mu B_{\infty}$$

Passing:

$$\mathcal{E} = \mathcal{E}_{\infty} + e\Phi_{\parallel} \longrightarrow \mathcal{E}_{\infty} = \mathcal{E} - e\Phi_{\parallel}$$



$$\mathcal{E}_{\parallel \infty} = \mathcal{E} - e\Phi_{\parallel} - \mu B_{\infty} = 0$$

Vlasov:

$$\frac{df}{dt} = 0$$

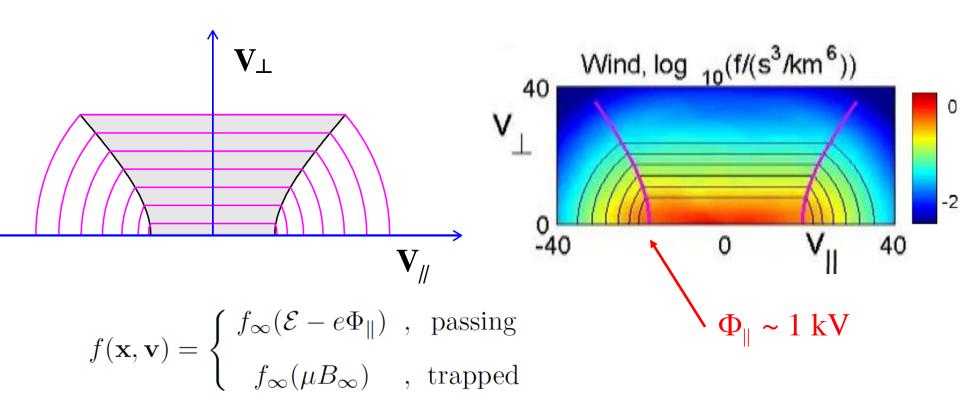
$$f(\mathbf{x}, \mathbf{v}) = f_{\infty}(\mathcal{E}_{\infty})$$

$$\Phi_{\parallel}(\mathbf{x}) = \int_{\mathbf{x}}^{\infty} \mathbf{E} \cdot d\mathbf{l}$$

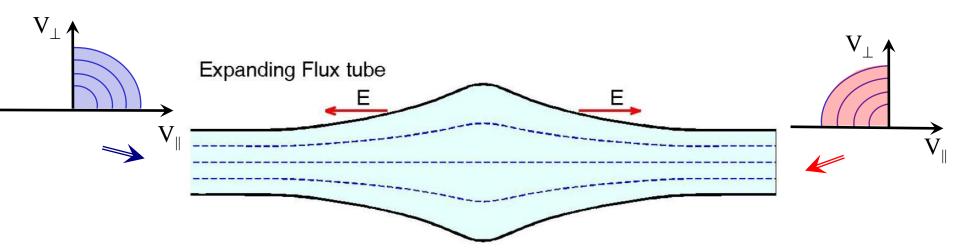
$$\int f_{\infty}(\mathcal{E} - e\Phi_{\parallel}) \quad \text{, passing}$$

$$f(\mathbf{x}, \mathbf{v}) = \begin{cases} f_{\infty}(\mathcal{E} - e\Phi_{\parallel}) &, \text{ passing} \\ f_{\infty}(\mu B_{\infty}) &, \text{ trapped} \end{cases}$$

Wind Spacecraft Observations in Distant Magnetotail, 60R_E



Formal derivation using an "ordering"



The drift kinetic equation:

$$\frac{\partial f}{\partial t} + (\mathbf{v}_{\parallel} + \mathbf{v}_{D}) \cdot \nabla f + \left[\mu \frac{\partial B}{\partial t} + e(\mathbf{v}_{\parallel} + \mathbf{v}_{D}) \cdot \mathbf{E} \right] \frac{\partial f}{\partial \mathcal{E}} = 0$$

Boundary conditions:

$$B = B_{\infty}$$

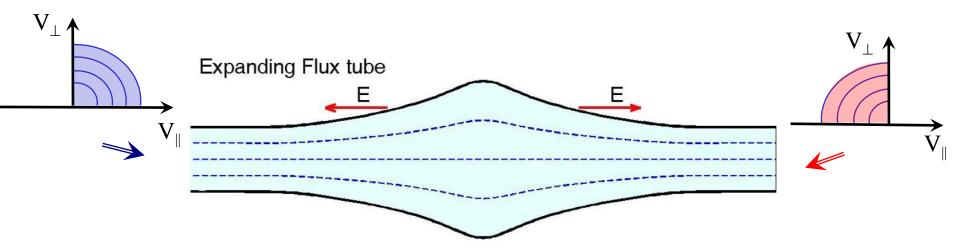
$$f = f_{\infty}(\mathcal{E}_{\parallel \infty}, \mathcal{E}_{\perp \infty})$$

Ordering:

$$abla_{\parallel} \sim \frac{1}{L} \quad , \qquad
abla_{\perp} \sim \frac{1}{d} \quad , \qquad \frac{\partial}{\partial t} \sim \frac{v_D}{d}$$

$$\frac{d}{L} \sim \delta \quad , \qquad \frac{v_D}{v_t} \sim \delta^2 \quad , \qquad \frac{E_{\parallel}}{E_{\perp}} \sim \delta$$

Formal derivation, passing electrons



Passing electrons, lowest order equation:

$$\mathbf{v}_{\parallel} \cdot \nabla f + e(\mathbf{v}_{\parallel} \cdot \mathbf{E}) \frac{\partial f}{\partial \mathcal{E}} = 0$$

Integrate along characteristics (field lines):

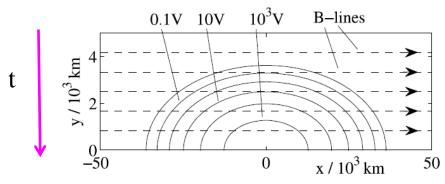
$$f(\mathcal{E}_{\parallel}, \mathcal{E}_{\perp}) = f_{\infty}(\mathcal{E} - e\Phi_{\parallel}^{\pm} - \mu B_{\infty}, \mu B_{\infty})$$
where
$$\mathcal{E} = \mathcal{E}_{\parallel} + \mathcal{E}_{\perp} \quad , \quad \Phi_{\parallel}^{\pm} = \int_{r}^{\pm \infty} \mathbf{E} \cdot d\mathbf{l}_{\parallel}$$

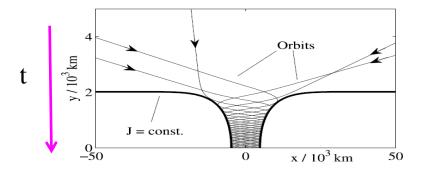
Quasi-neutrality at boundaries: $n = \int f d^3v = n_{\infty}$ (only passing electrons here)

Formal derivation, trapped electrons

Use full equation with solution: $f = g(\mu, J, \phi_J)$ = $f_{\infty} \left(\mathcal{E}_{\parallel \infty} [\mu, J, \phi_j], \mathcal{E}_{\perp \infty} [\mu, J, \phi_j] \right)$

 2^{nd} adiabatic invariant: $J = \oint v_{\parallel} dl$





$$\mathcal{E}_{\perp \infty} = h_1(\mu, J, \phi_j) = \mu B_{\infty}$$

$$\mathcal{E}_{\parallel \infty} = h_2(\mu, J, \phi_j) \approx 0$$

Only electrons with small parallel energy will be caught in the magnetic and electric well as it develops slowly compared to the electron transit time, τ ($\sim L/v_t$).

Thus, for Maxwellian
$$f_{\infty}$$
:

$$f(\mathbf{x}, \mathbf{v}) = \begin{cases} f_{\infty}(\mathcal{E} - e\Phi_{\parallel}) &, \text{ passing} \\ f_{\infty}(\mu B_{\infty}) &, \text{ trapped} \end{cases}$$

$$\mathcal{E}_{\parallel \infty} \leq \mu \frac{\partial B}{\partial t} \tau + e \frac{\partial E_{\parallel}}{\partial t} \tau L$$

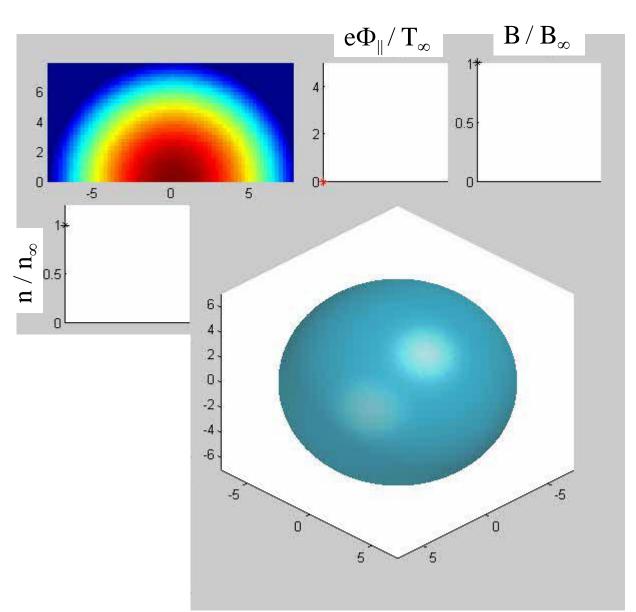
$$\simeq \mu B \frac{v_d}{d} \frac{L}{v_t} + e E_{\parallel} \frac{v_d}{d} \frac{L}{v_t} L$$

$$\simeq \delta \left(T_e + e \Phi_{\parallel} \right) .$$

Kinetic Model Fluid Closure

Theoretical distribution:

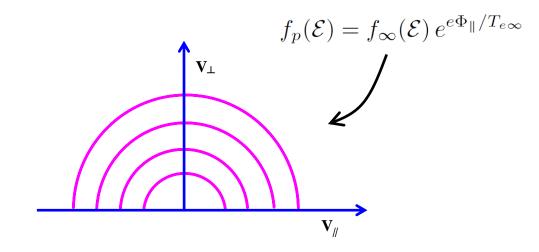
$$f(\mathbf{x}, \mathbf{v}) = \begin{cases} f_{\infty}(\mathcal{E} - e\Phi_{\parallel}) &, \text{ passing} \\ f_{\infty}(\mu B_{\infty}) &, \text{ trapped} \end{cases}$$



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Kinetic Model Fluid Closure



$$\Phi_{\parallel} < 0, B > B_{\infty} \implies \text{no trapping}$$

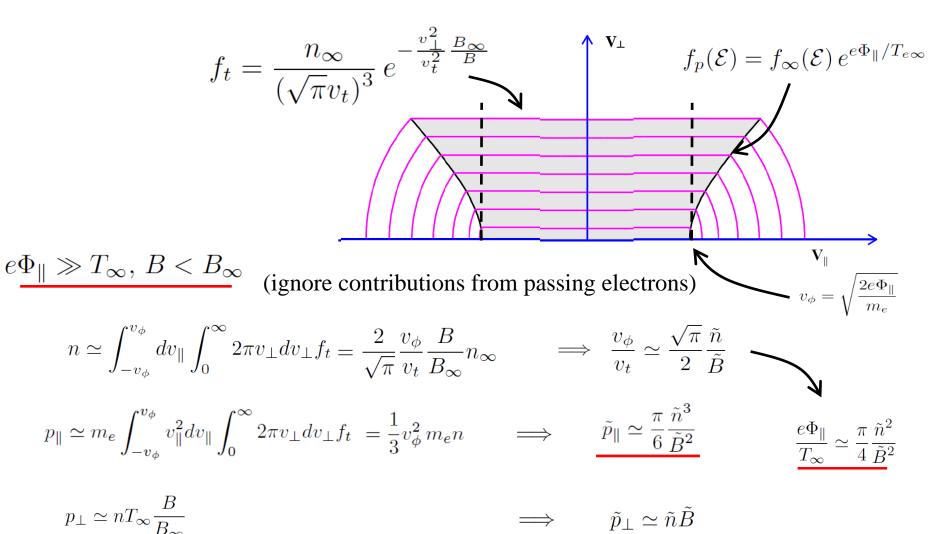
$$n = n_{\infty} e^{e\Phi_{\parallel}/T_{\infty}} < n_{\infty} , \qquad \frac{e\Phi_{\parallel}}{T_{\infty}} = \log\left(\frac{n}{n_{\infty}}\right)$$

$$\frac{e\Phi_{\parallel}}{T_{\infty}} = \log\left(\frac{n}{n_{\infty}}\right)$$

$$p_{\parallel} = p_{\perp} = nT_{\infty}$$

$$\tilde{p}_{\parallel} = \tilde{p}_{\perp} = \tilde{n}$$

Kinetic Model Fluid Closure



Kinetic Model -> Fluid Closure (EoS)

$$f(\mathbf{x},\mathbf{v}) = \begin{cases} f_{\infty}(\mathcal{E} - e\Phi_{\parallel}) &, \text{ passing} \\ f_{\infty}(\mu B_{\infty}) &, \text{ trapped} \end{cases}$$

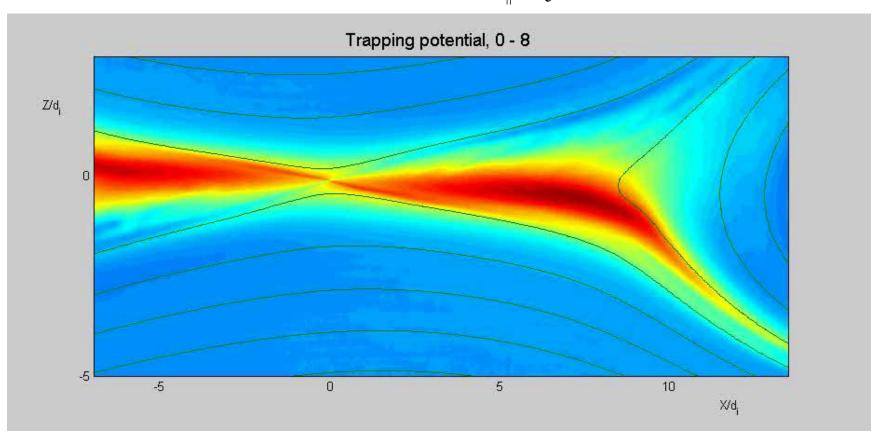
$$\int \dots \, \mathbf{d}^{3}\mathbf{v} \qquad \qquad \Phi_{\parallel} = \Phi_{\parallel}(n,B) \qquad \qquad$$



A. Le et al., PRL (2009)

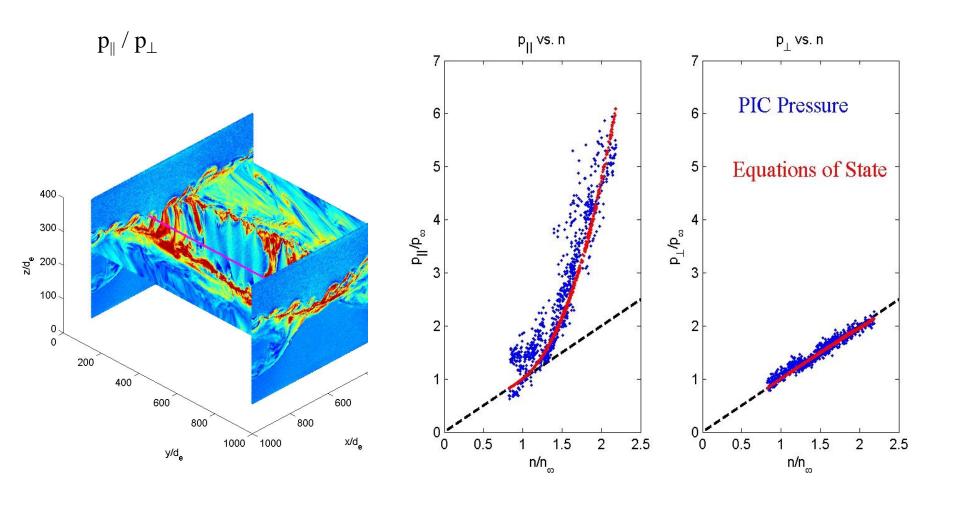
The Acceleration Potential in a Kinetic Simulation

 $e\Phi_{\parallel}/T_{e}$



Confirmed in Kinetic Simulations

EoS previously confirmed in 2D simulations, now also in 3D simulations.



New EoS Now Implemented in Two-Fluid Code

New code implemented by O Ohia using the HiFi framework developed in part by VS Lukin

Standard two-fluid equations

$$\begin{split} \frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{V}_i) &= 0 \\ m_i n \left(\frac{\partial \mathbf{V}_i}{\partial t} + \mathbf{V}_i \cdot \nabla \mathbf{V}_i \right) &= \mathbf{J} \times \mathbf{B} - \nabla \cdot \bar{\mathbf{P}} + m_i n \nu_i \nabla^2 \mathbf{V}_i \\ \frac{\partial}{\partial t} \left(\frac{p_i}{n^{\Gamma}} \right) &= -\mathbf{V}_i \cdot \nabla \frac{p_i}{n^{\Gamma}} \\ \frac{\partial \mathbf{B}'}{\partial t} &= -\nabla \times \mathbf{E}' \\ \mathbf{E}' + \mathbf{V}_i \times \mathbf{B} &= \frac{1}{ne} \left(\mathbf{J} \times \mathbf{B}' - \nabla \cdot \bar{\mathbf{P}}_e \right) + \eta_R \mathbf{J} - \eta_H \nabla^2 \mathbf{J} \\ \mathbf{B}' &= \left(1 - d_e^2 \nabla^2 \right) \mathbf{B} \\ \mu_0 \mathbf{J} &= \nabla \times \mathbf{B} \end{split}$$

Anisotropic pressure model

$$\bar{\mathbf{P}} = p_i \bar{\mathbf{I}} + \bar{\mathbf{P}}_e = p_i \bar{\mathbf{I}} + p_\perp \bar{\mathbf{I}} + \left(p_{\parallel} - p_\perp\right) \frac{\mathbf{B}\mathbf{B}}{B^2}$$

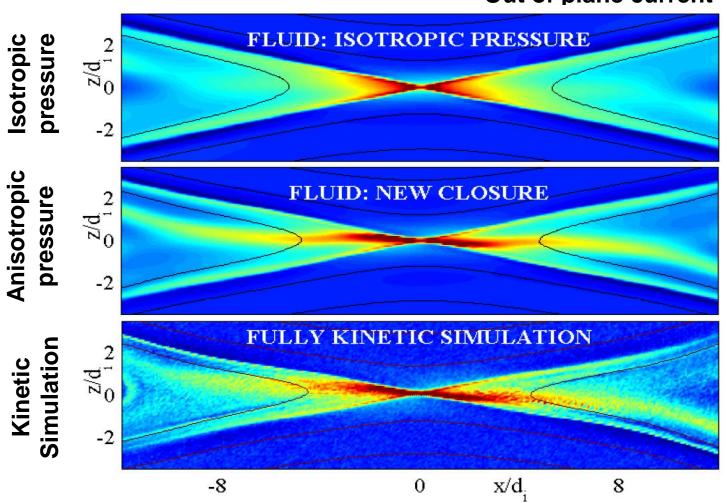
$$p_{*\parallel} = \frac{n_*}{1 + \alpha/2} + \frac{\alpha \pi}{6 + 3/\alpha}$$

$$p_{*\perp} = \frac{n_*}{1 + \alpha} + \frac{n_* B_*}{1 + 1/\alpha}$$

where $\alpha = n_*^3/B_*^2$ and for any quantity $Q, Q_* = Q/Q_{\infty}$

New *EoS* Implemented in Two-Fluid Code

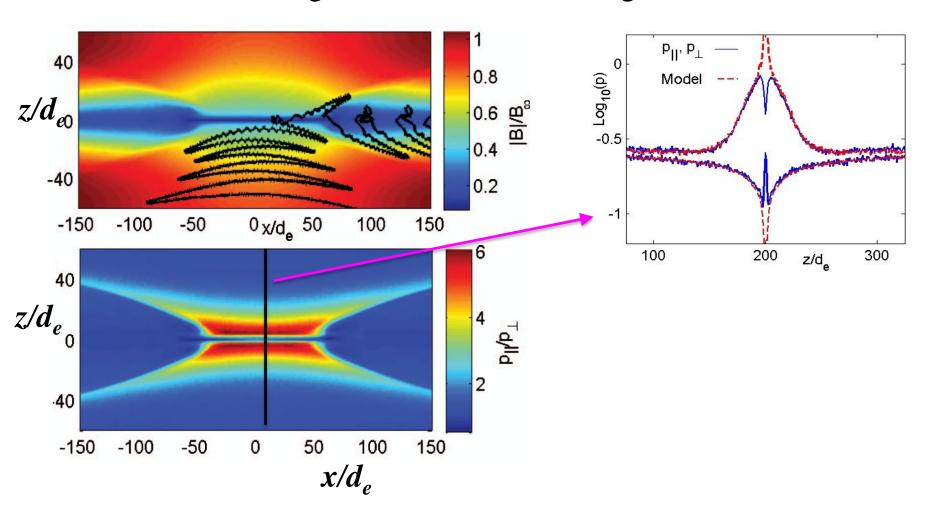




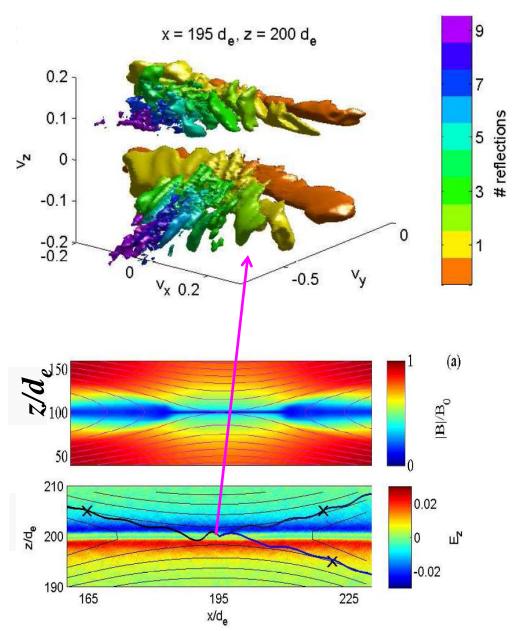
Ohia et al., PRL 2012 (in press)

EoS for anti-parallel reconnection?

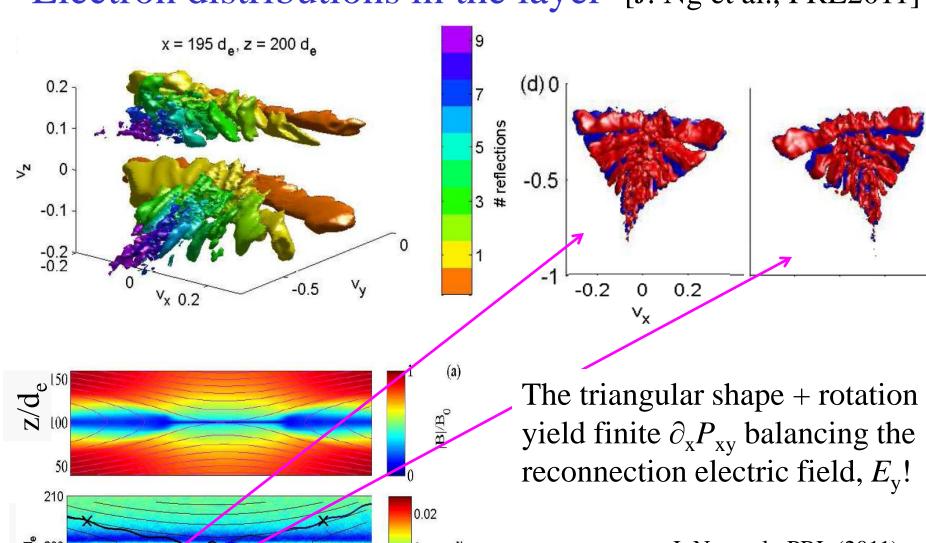
The electrons are magnetized in the inflow region:



Electron distributions in the layer [J. Ng et al., PRL2011]



Electron distributions in the layer [J. Ng et al., PRL2011]



-0.02

225

195

x/de

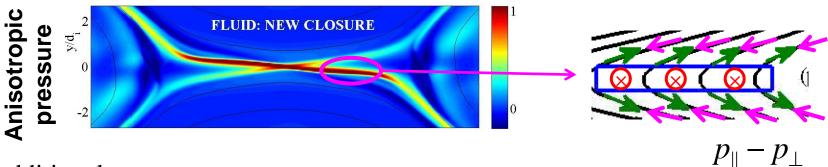
165

J. Ng et al., PRL (2011)

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Analytic Model for Electron Jets



Additional current term:

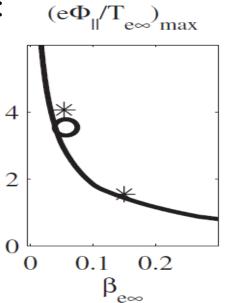
$$J_{\perp \mathrm{extra}} = [(p_{\parallel} \! - \! p_{\perp})/B] \hat{b} \! \times \! \hat{b} \! \cdot \! \nabla \hat{b}$$

The magnetic tension is balanced by pressure anisotropy:

$$p_{\parallel}(n,B) - p_{\perp}(n,B) = B^2 / \mu_0$$

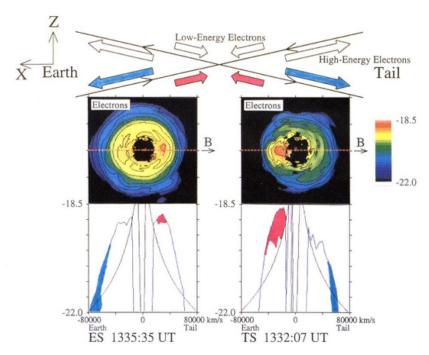
Use *EoS* to get scaling laws:

$$\beta_{e\infty} = \frac{\text{plasma pressure}}{\text{magnetic pressure}}$$



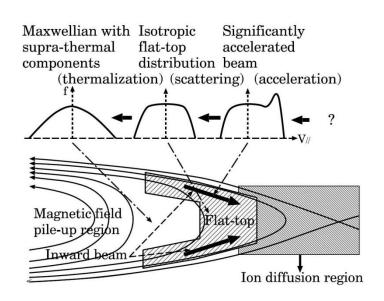
Observed Electron Heating

1-10 keV beam like electron distributions often observed



T. Nagai, et al., *J. Geophys. Res.*, **106**, 25.929 (2001).

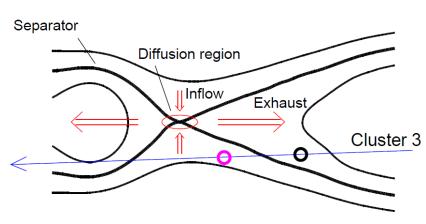
Flat-top distributions typical in the exhaust; related beams?

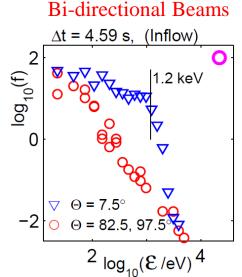


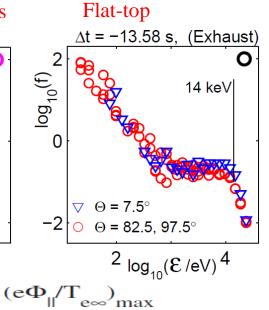
Y. Asano, et al., *J. Geophys. Res.*, *113*(A1), (2008).

Observed Electron Heating

Cluster observations on 2001-10-01.





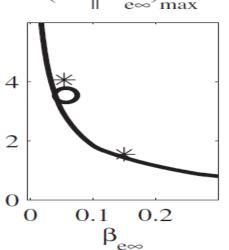


Lobe: $T_e \sim 0.1 \text{ keV}$

Inflow: $T_{e\perp} \sim 0.1 \text{ keV}$, $T_{e\parallel} \sim 1 \text{ keV}$

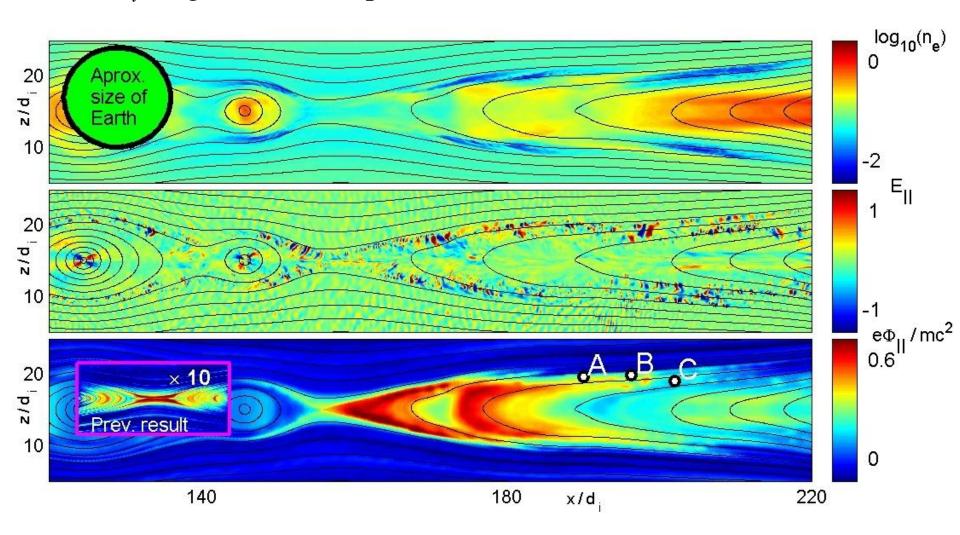
Exhaust: $T_{e\perp} \sim T_{e\parallel} \sim 10 \text{keV}$

For this event $\beta_e \sim 0.003$



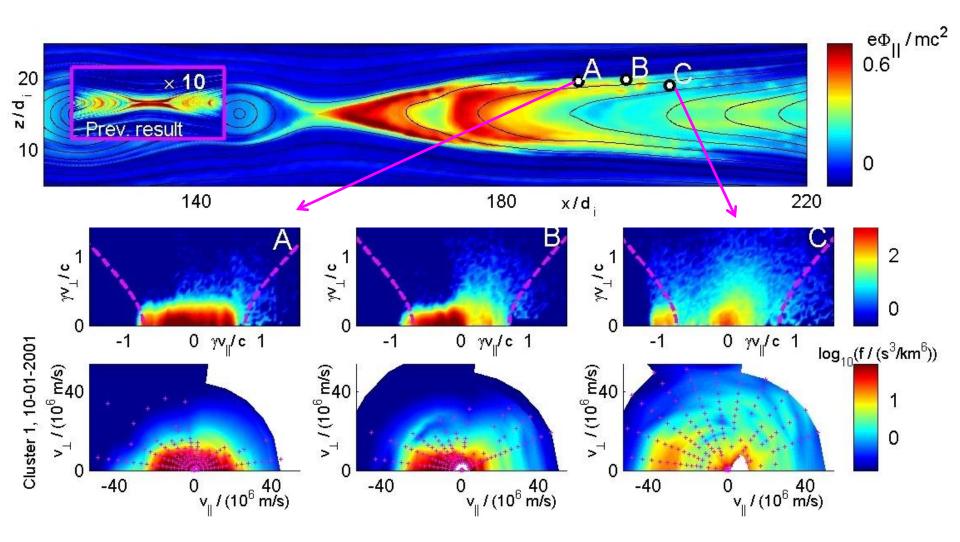
New simulation with $\beta_e \sim 0.003$

• 320 d_i long, 180 billion particles!



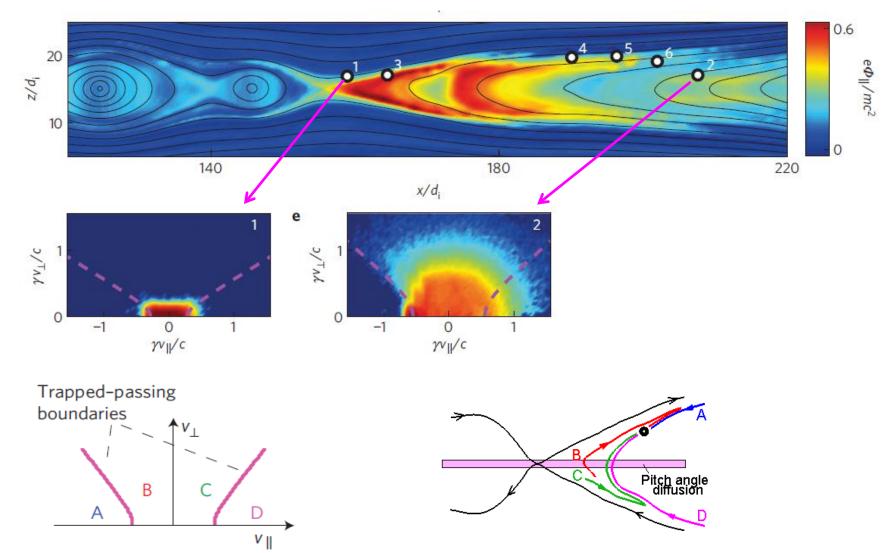
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• 320 d_i long, 180 billion particles!



Flat-top Distributions

• 320 d_i long, 180 billion particles!

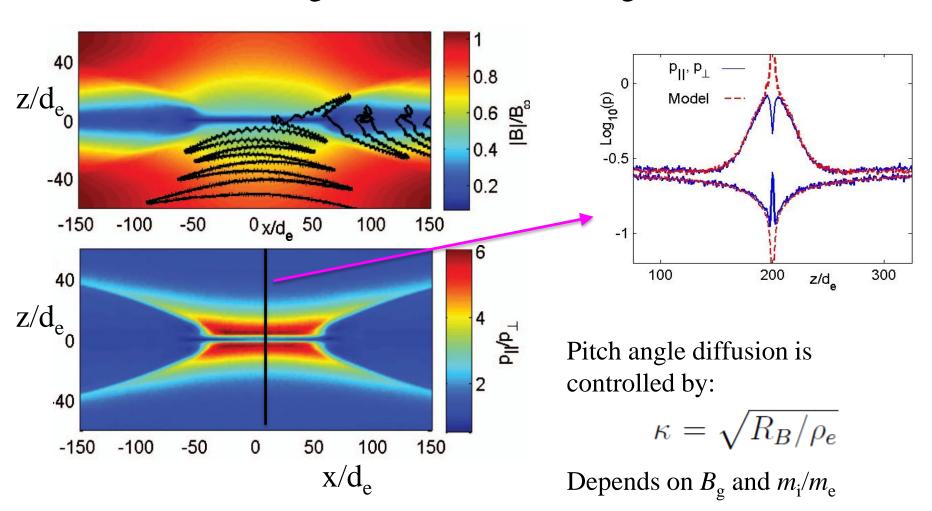


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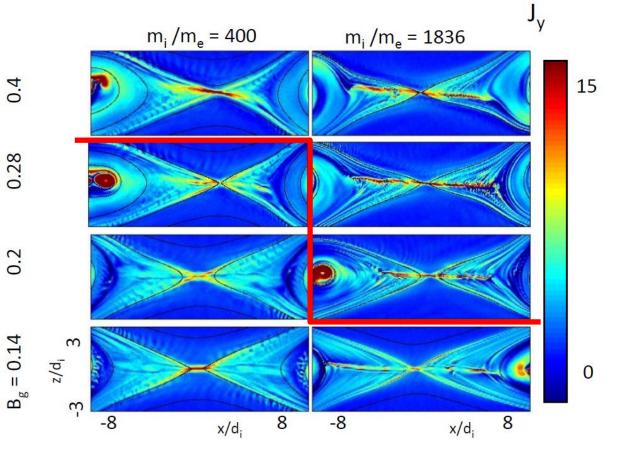
EoS for Anti-Parallel Reconnection?

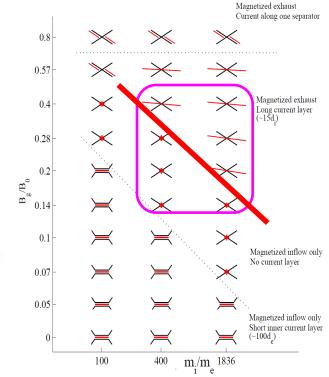
The electrons are magnetized in the inflow region:

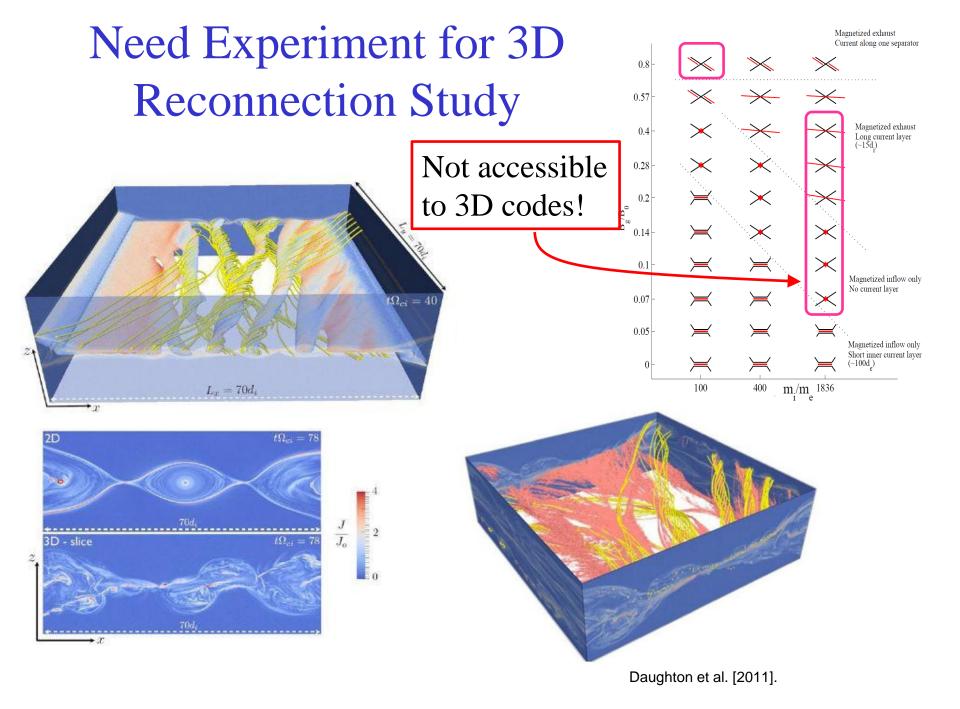


Scan in $B_{\rm g}$ and $m_{\rm i}/m_{\rm e}$ $(\beta_{\rm e} = 0.03)$ Magnetized exhaust Current along one separator 0.8 0.57 Magnetized exhaust 0.4 Long current layer (~15d_i) Pitch angle diffusion 0.28 0.2 0.14 0.1 Magnetized inflow only No current layer 0.07 0.05 Magnetized inflow only Short inner current layer (~100d_e) 0 400 100

Scan in $B_{\rm g}$ and $m_{\rm i}/m_{\rm e}$







Requirements on New Experiment

- Large normalized size of experiment: $L/d_i \sim 10$ (high *n*, large *L*)
- Low collisionallity to allow $p_{\parallel} >> p_{\perp}$: $\tau_{\rm ei} \ v_{\rm A} > d_{\rm i}$ (low n, high $T_{\rm e}$, high B)
- Low electron pressure: $\beta_{\rm e} < 0.05$ (low $n, T_{\rm e}$, high B)
- Manageable loop voltage: $0.1v_A B_{rec} (2\pi R) < 5kV$ (high n, low B)
- Variable guide field: $B_g = 0 4 B_{rec}$
- Symmetric inflows

Experimental window available in Hydrogen or Helium plasma with

$$n \sim 10^{18} \text{ m}^{-3}$$
, $T_{\rm e} \sim 15 \text{ eV}$, $B_{\rm rec} \sim 15 \text{ mT}$, $L \sim 2 \text{ m}$

Flare heating by parallel E-fields?

Ohm's law:

$$-enE_{\parallel} = \hat{\mathbf{b}} \cdot (\nabla \cdot \mathbf{p})$$

Before reconnection: $p = nT_e \rightarrow e\Phi_{\parallel} \sim T_e \log(n/n_0)$

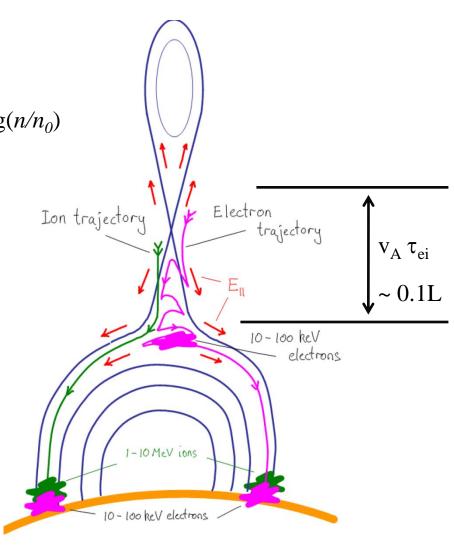
During reconnection: $p_{\parallel} \propto \frac{n^3}{B^2}$

$$\bullet \qquad e\Phi_{\parallel} \approx T_e \frac{(n/n_0)^2}{(B/B_0)^2}$$

$$(n/n_0) \approx 10$$
, $(B/B_0) \approx 0.5$

$$e\Phi_{\parallel} \approx 400T_e$$





Conclusions

- A new analytic model for electron the electron distribution function was inspired by the VTF experiment and has been confirmed in kinetic simulations.
- The model has been applied as a closure to the fluid equations and has helped explain electron energization in spacecraft observations.
- Long current layer can be driven by the pressure anisotropy for magnetotail conditions and their investigation requires a new experimental facility.